

# **Fast Approximate Broadband Phase Retrieval for Segmented Systems**

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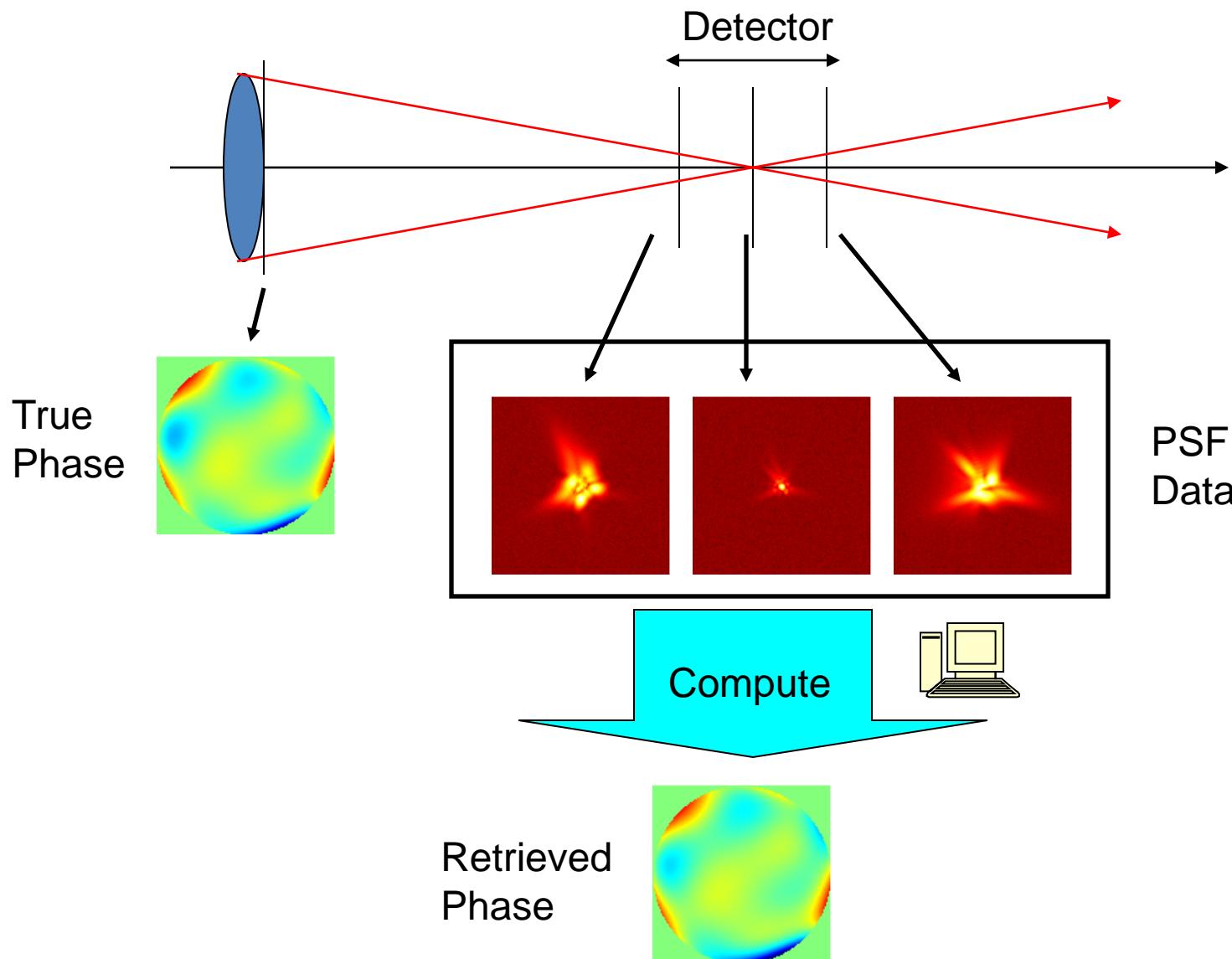
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# Outline

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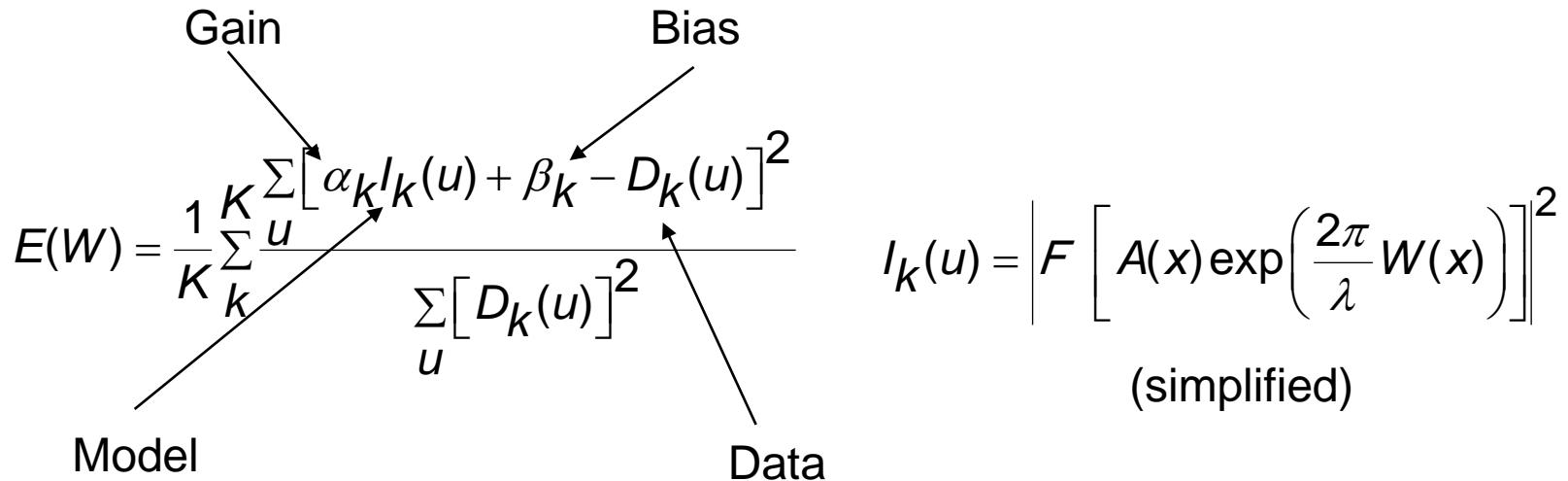
- Application: Wavefront sensing for large telescopes
  - Space telescopes (James Webb Space Telescope)
  - Adaptive Optics for ground based telescopes
- Wavefront sensing method: Focus diverse phase retrieval
- Case of interest: Extremely broadband sources
- Problem
  - Monochromatic phase retrieval algorithms fail above  $\sim 10\%$  fractional bandwidth
  - Broadband (polychromatic) algorithms are computationally expensive
- Solution
  - Employ monochromatic algorithm with approximation method
- Result
  - 270x speed up in computational performance for 133% fractional bandwidth
  - Acceptable accuracy
  - Accuracy better for monolithic systems, worse for segmented
- Investigate difference in accuracy between segmented and monolithic systems

# Focus Diverse Phase Retrieval



- Given
  - System parameters (F/#, detector pixel pitch, etc.)
  - Measured point spread function images (e.g. images of stars) through focus
  - Clear aperture of pupil
- Determine
  - Wavefront error in the exit pupil of the system (phase)
- Issues
  - Optical fields are complex-valued quantities
  - Measured data (intensity) has only magnitude, no phase
  - Measured data corrupted with noise
- Solution strategies
  - Gerchberg-Saxton / iterative transform type algorithms
  - Non-linear optimization type algorithms

- Parameterize problem in terms of a set of variables (e.g. Zernike coefficients)
- Generate wavefront ( $W$ ) and/or field from parameters
- Propagate field to PSF plane
- Use an error metric to evaluate agreement between model and data<sup>1</sup>
- Use standard non-linear optimization algorithm to reduce error metric value
- Analytic expressions for error metric gradients allow use of gradient search algorithms



The diagram illustrates the error metric  $E(W)$  as a function of wavefront parameters  $W$ . The metric is given by:

$$E(W) = \frac{1}{K} \sum_k \frac{u}{K} \left[ \alpha_k I_k(u) + \beta_k - D_k(u) \right]^2 + \sum_u \left[ D_k(u) \right]^2$$

Annotations with arrows point to the components:

- Gain** points to the term  $\alpha_k$ .
- Bias** points to the term  $\beta_k$ .
- Model** points to the term  $I_k(u)$ .
- Data** points to the term  $D_k(u)$ .

Below the diagram, the field intensity  $I_k(u)$  is defined as:

$$I_k(u) = \left| F \left[ A(x) \exp \left( \frac{2\pi}{\lambda} W(x) \right) \right] \right|^2$$

(simplified)

<sup>1</sup>S. T. Thurman and J. R. Fienup, "Phase retrieval with signal bias," *J. Opt. Soc. Am. A* 26, 1008–1014 (2009).

- Use cases for broadband phase retrieval
  - Narrow spectral filters unavailable
  - Dim sources
  - Low throughput due to misalignment
  - Short exposures times
    - Pointing instability (space)
    - Atmospheric instability (ground based AO)
  - Segment piston determination
- Traditional approach<sup>1,2</sup>
  - Simulate multiple individual wavelengths
  - Rule of thumb: 1 wavelength per 5% fractional bandwidth
  - Add incoherently to simulate polychromatic PSF
  - Do normal nonlinear optimization phase retrieval

<sup>1</sup> J. R. Fienup, "Phase retrieval for undersampled broadband images," *J. Opt. Soc. Am. A* **16**, 1831–1837 (1999).

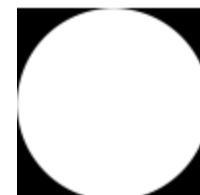
<sup>2</sup> G. R. Brady and J. R. Fienup, "Effect of broadband illumination on reconstruction error of phase retrieval in optical metrology," *Proc. SPIE* 6617, 661701 (2007).

- Fine for modest bandwidth
- Very computationally intensive for large bandwidths
  - More wavelengths → more computation
    - Linear in number of wavelengths
  - Shorter wavelengths → larger arrays → more computation
    - Pupil sampling requirement
      - Avoid  $\pi$  phase jumps
    - Phase for given OPD increases with shorter wavelength
    - Array size inversely proportional to wavelength
    - Cost approximately quadratic in array size

Monochromatic Pupil  
3  $\mu\text{m}$ , 128x128



Broadband Pupil  
1-5  $\mu\text{m}$ , 384x384



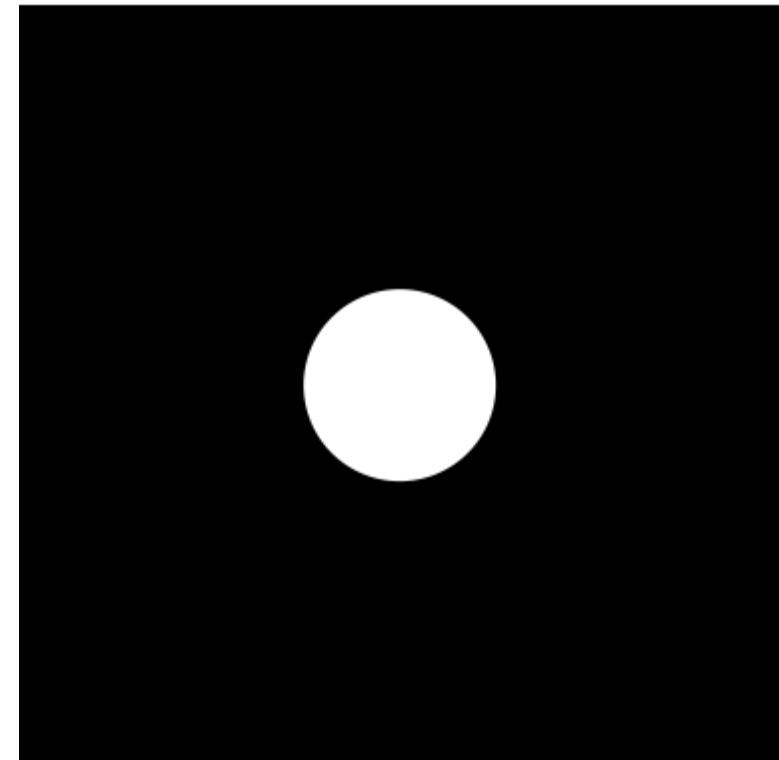
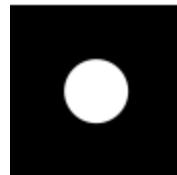
# Array Sizes

- Shorter / longer wavelengths → under / over sampling  
→ more computation

- PSF initially computed on large array
- Requires padding in pupil domain

Broadband Pupil  
(largest) 1518x1518

Monochromatic Pupil  
342x342



- Achromatic system
  - Reflecting telescopes
  - Color corrected instruments
- OPD / rays the same at all wavelengths
- Geometrical optics spot diagram the same at all wavelengths
- For poorly corrected systems geometrical optics should predict PSF shape
- PSF shape should be roughly the same at all wavelengths
- Only high frequency diffraction effects depend strongly on wavelength

# Approximation Procedure

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- Data  $D$ , blurring kernel  $B$ , forward model  $F$ , wavefront  $W$ ,  $K$  image planes
- Blur measured PSF data with Gaussian kernel (suppress diffraction effects)

$$D_{B,k} = B \otimes D_k$$

- Simulate monochromatic PSF at the center wavelength

$$I_k = F_{\lambda,k}[W]$$

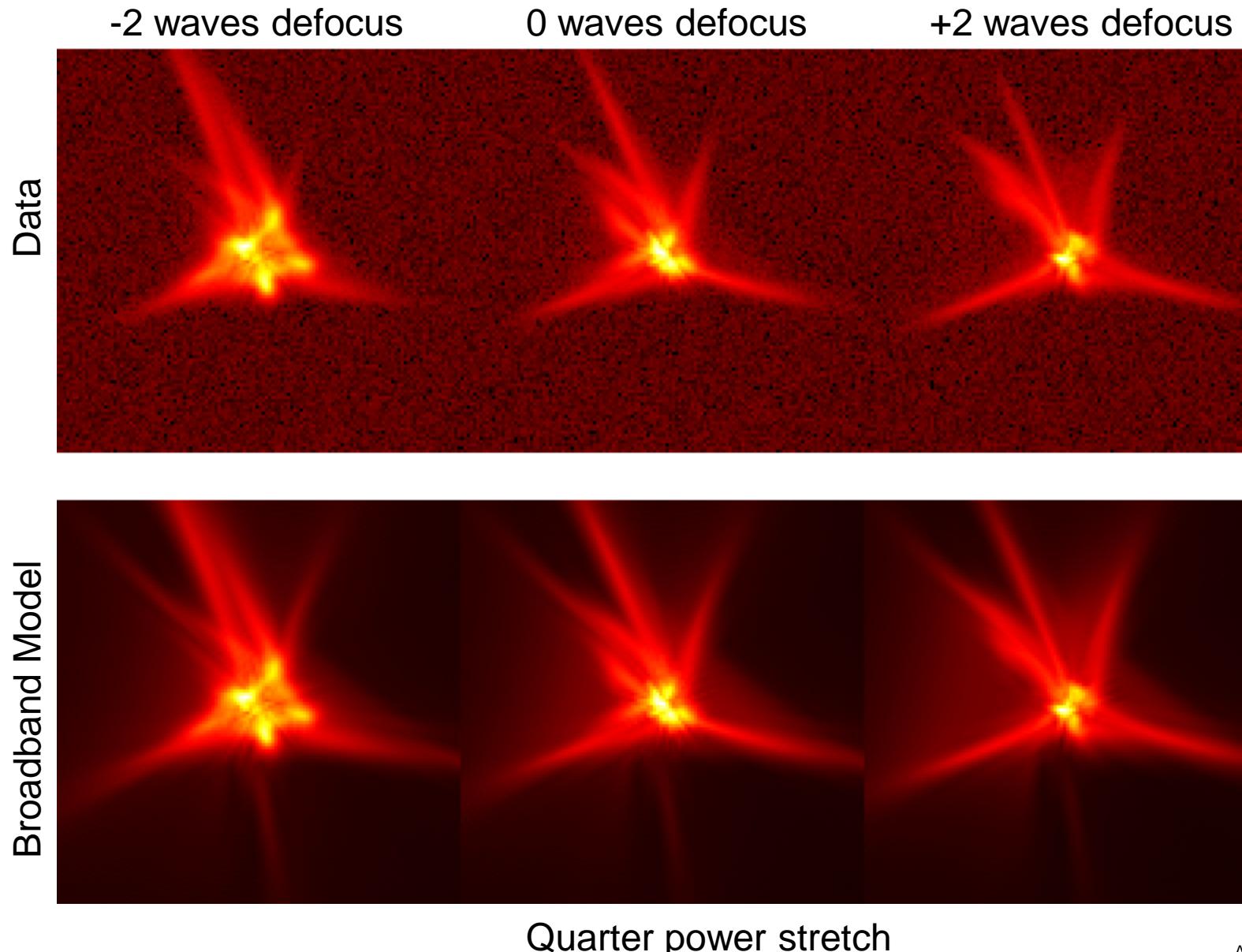
- Blur modeled PSF with same Gaussian kernel

$$I_{B,k} = B \otimes I_k = B \otimes F_{\lambda,k}[W]$$

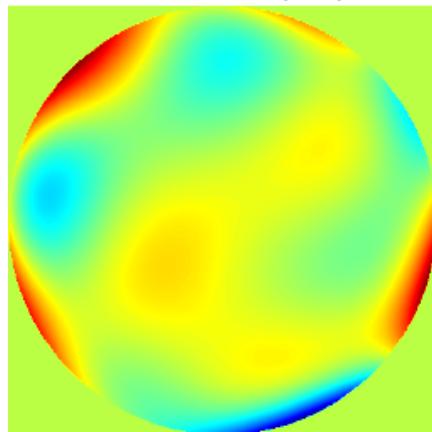
- Use non-linear optimization to fit blurred data against blurred model

$$E(W) = \frac{\frac{1}{K} \sum_k \frac{\sum_u \left[ \alpha_k I_{B,k}(u) + \beta_k - D_{B,k}(u) \right]^2}{\sum_u \left[ D_{B,k}(u) \right]^2}}{K}$$

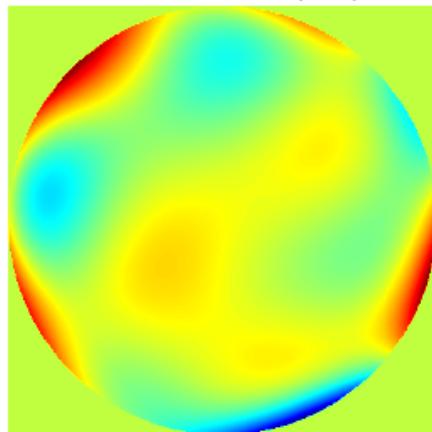
- Similar to JWST Fine Guidance Sensor (FGS)
- F/8 system
- 18  $\mu\text{m}$  pixels
- 1-5  $\mu\text{m}$  bandwidth (133% fractional bandwidth)
  - Flat spectrum
- Modeled detector area: 128x128 pixels (artificially small)
- Noise model: 50,000 photons in peak pixel, 25 photons of read noise, 100 photon noise-free bias
- Measured point spread functions at -2, 0, 2 waves center to edge defocus
- Modeled with monolithic aperture
- Monte Carlo simulation
  - 64 trials
  - Minimum wavefront error: 0.025 waves RMS
  - Maximum wavefront error: 0.500 waves RMS



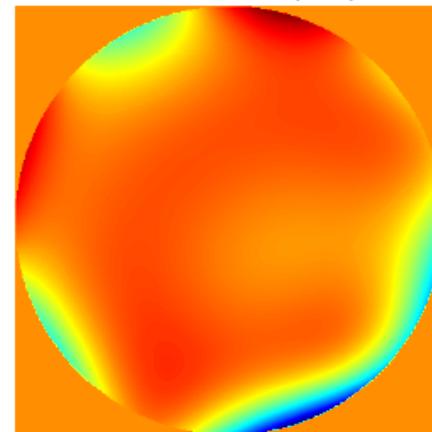
True OPD (wv)



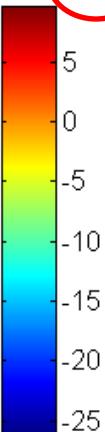
Result OPD (wv)



Error OPD (wv)



$\times 10^{-3}$



OPD in waves

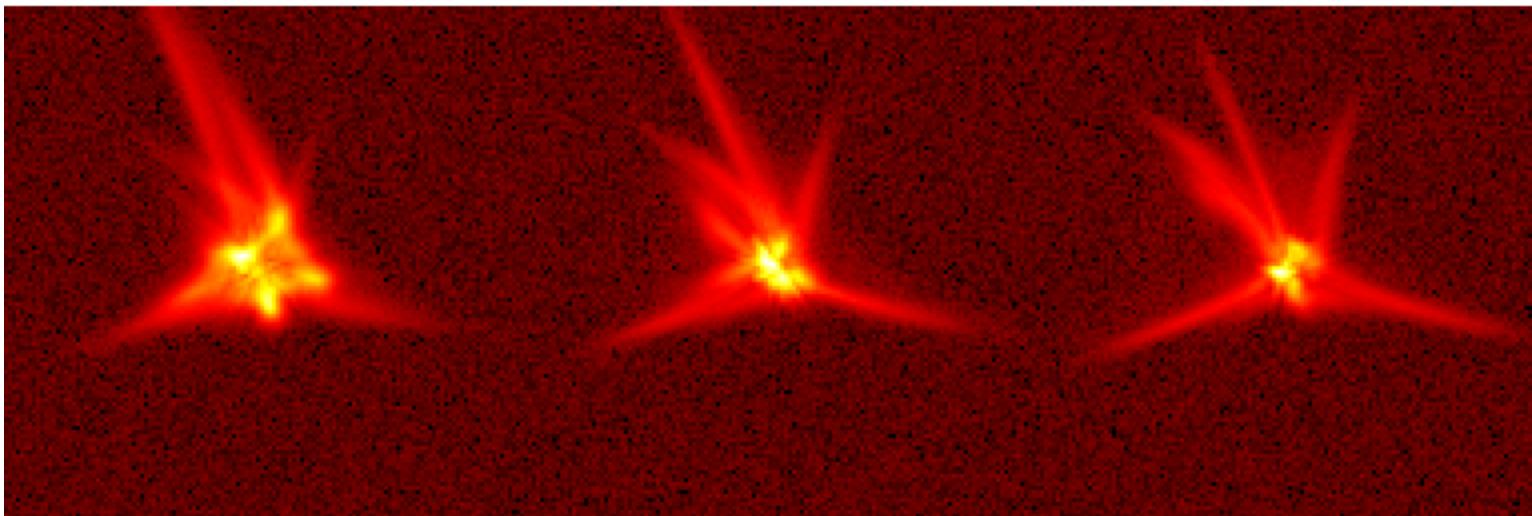
# Unblurred PSFs (Monolithic)

-2 waves defocus

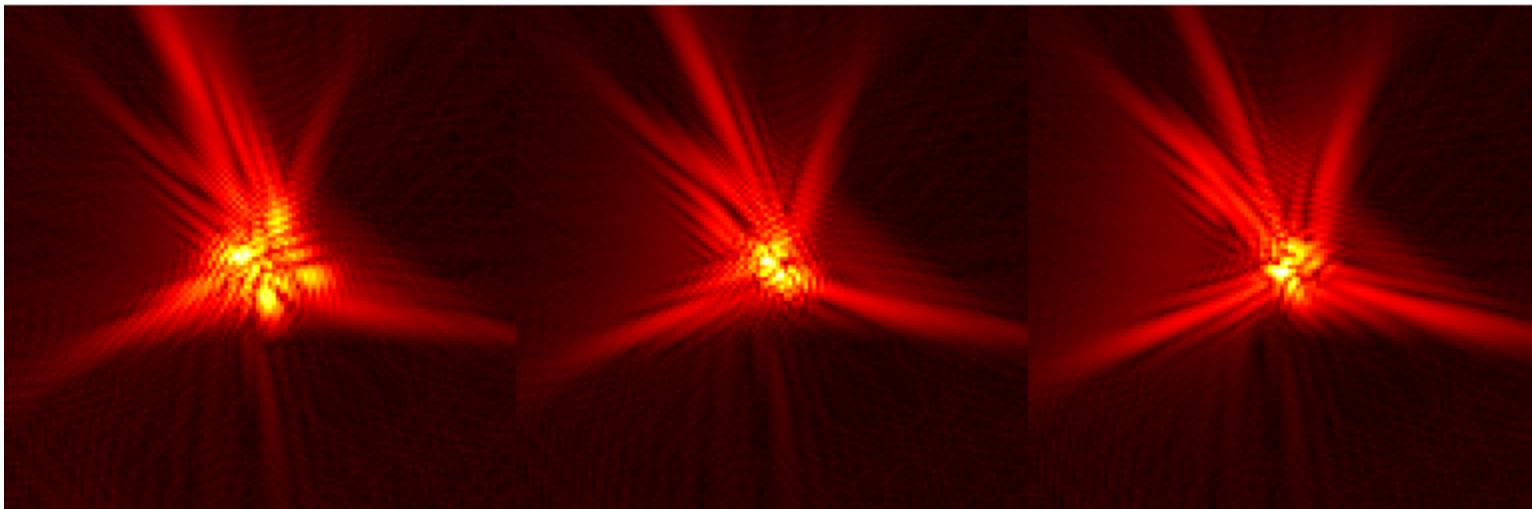
0 waves defocus

+2 waves defocus

Data



Unblurred  
Monochromatic



Quarter power stretch

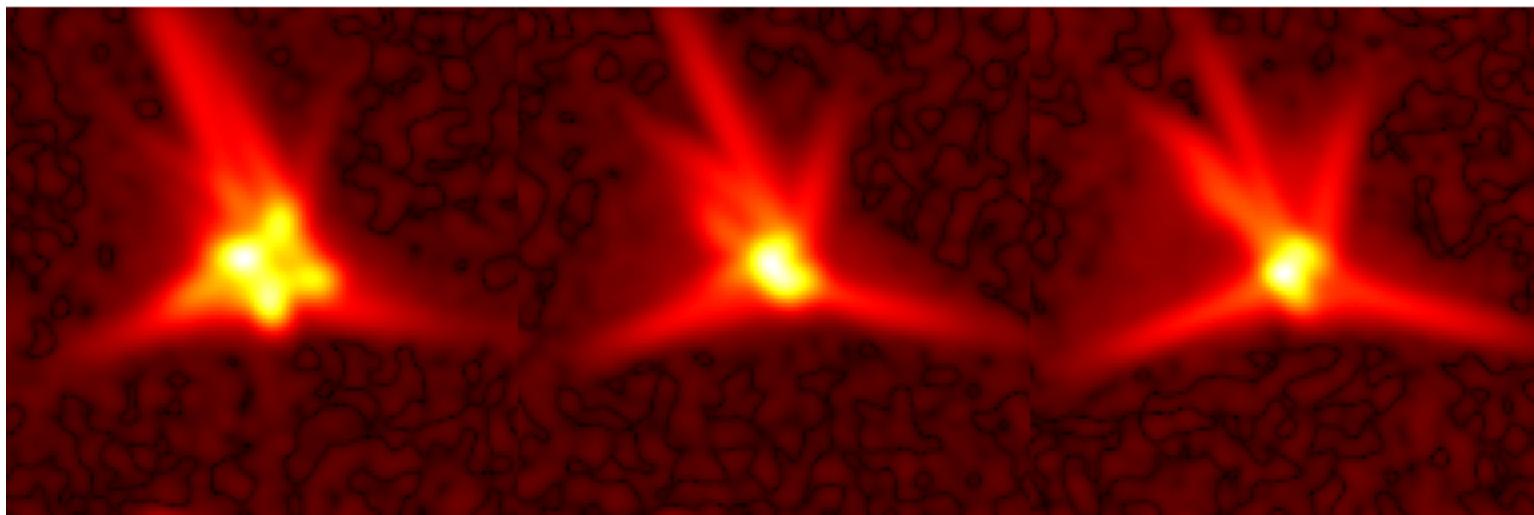
## Blurred PSFs (Monolithic)

-2 waves defocus

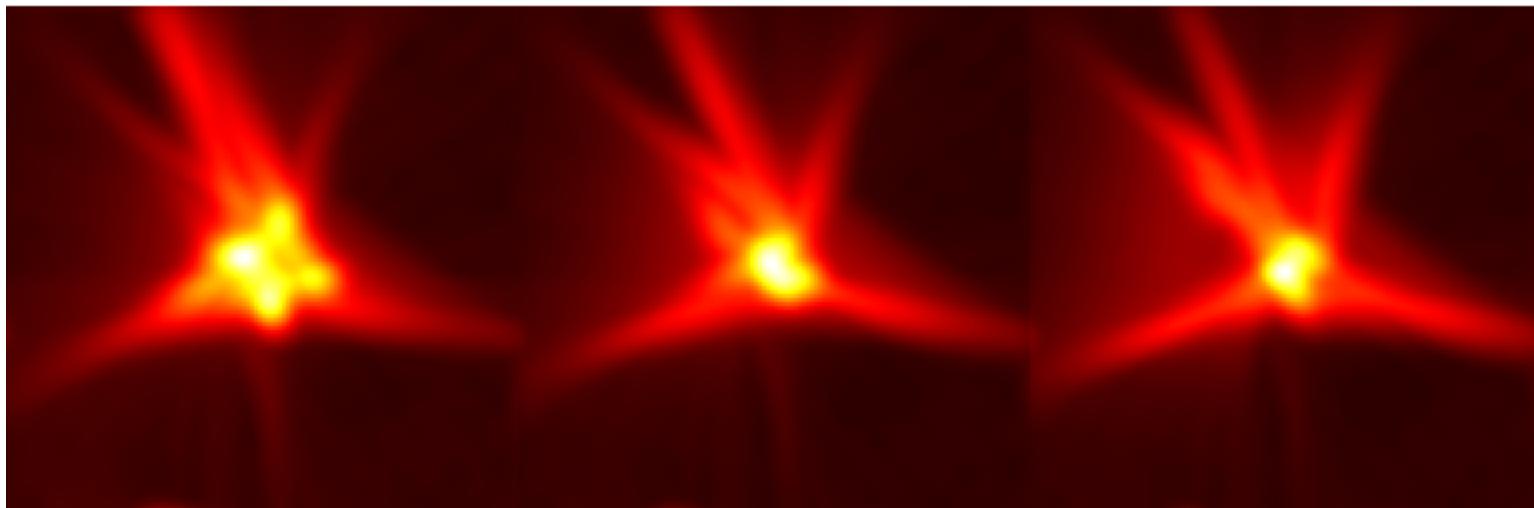
0 waves defocus

+2 waves defocus

Data

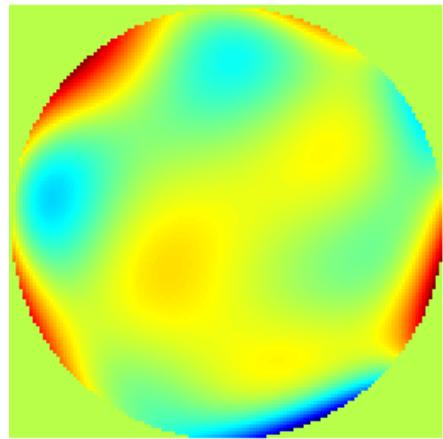


Blurred  
Monochromatic

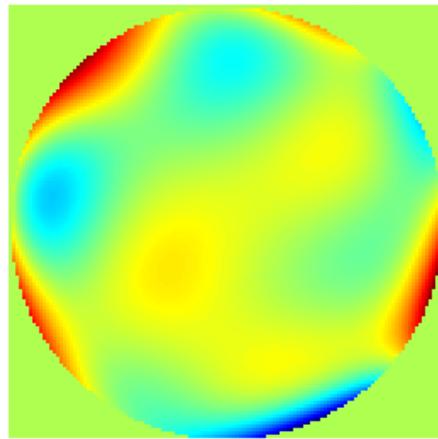


Quarter power stretch

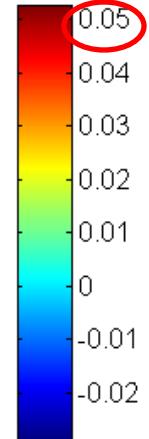
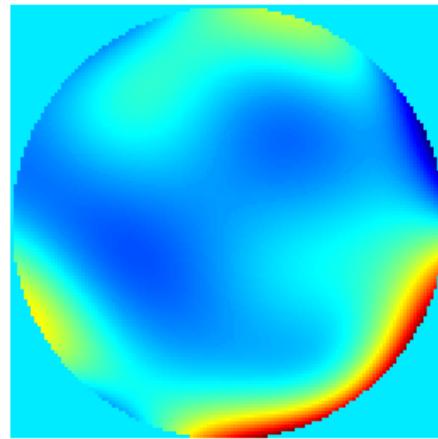
True OPD (wv)



Result OPD (wv)

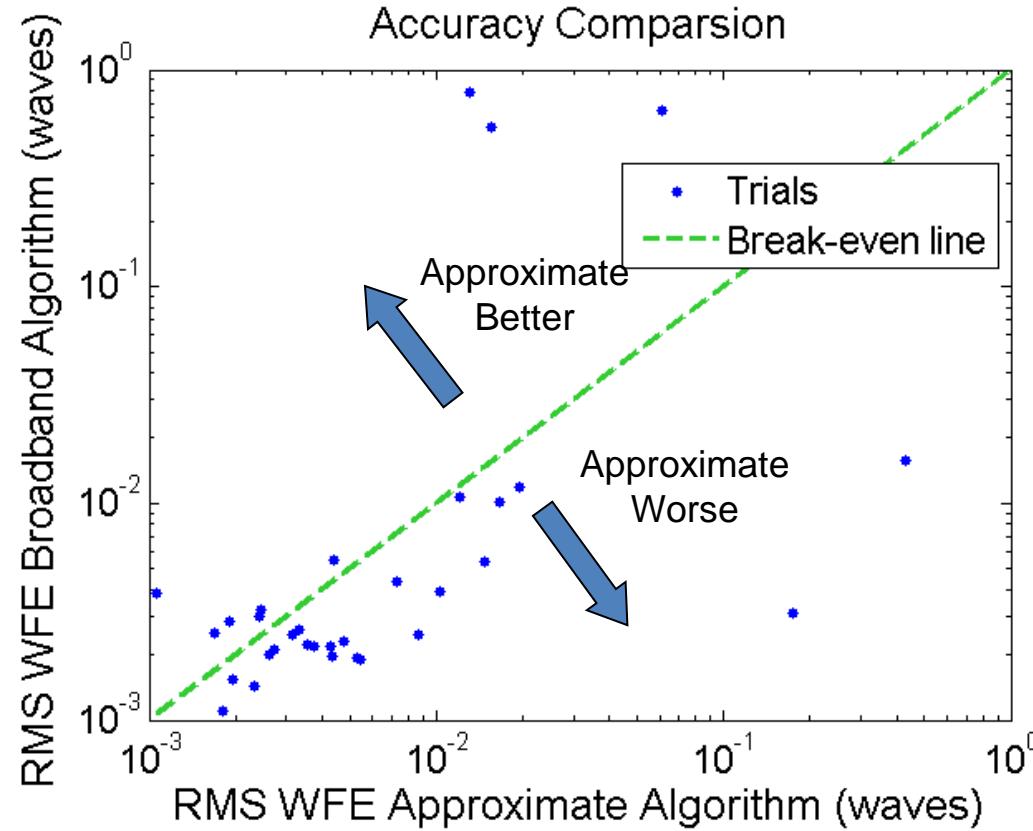


Error OPD (wv)

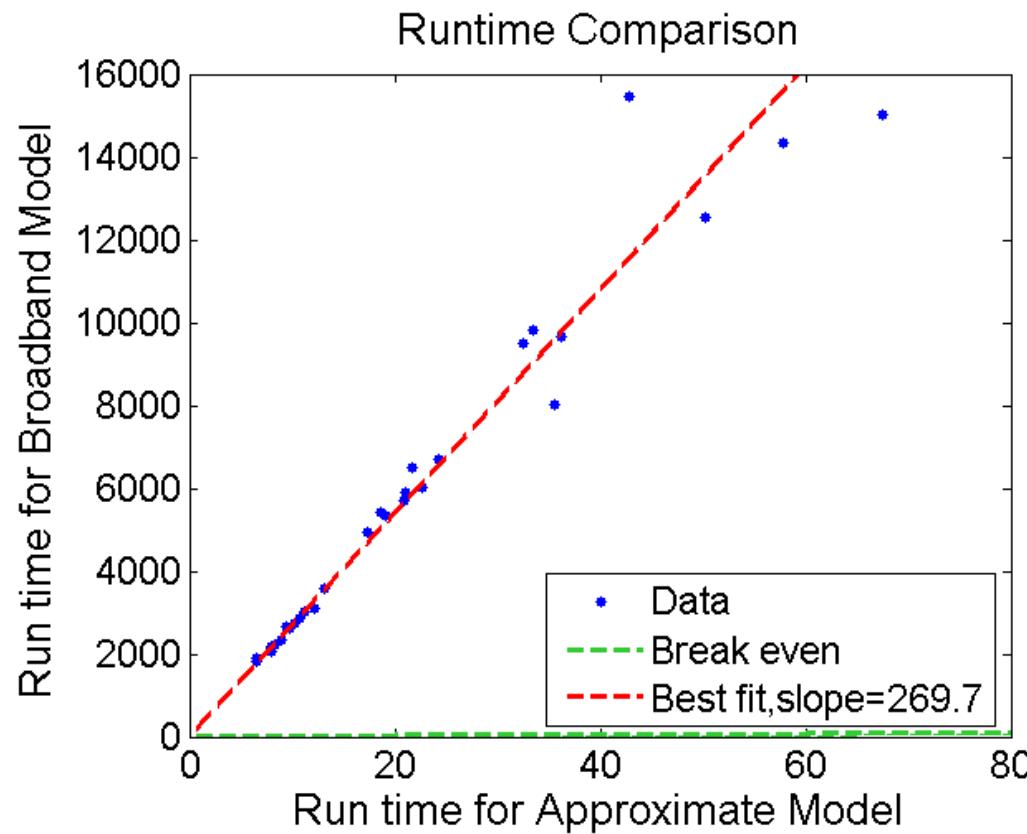


OPD in waves

# Monte Carlo Results



# Monte Carlo Results



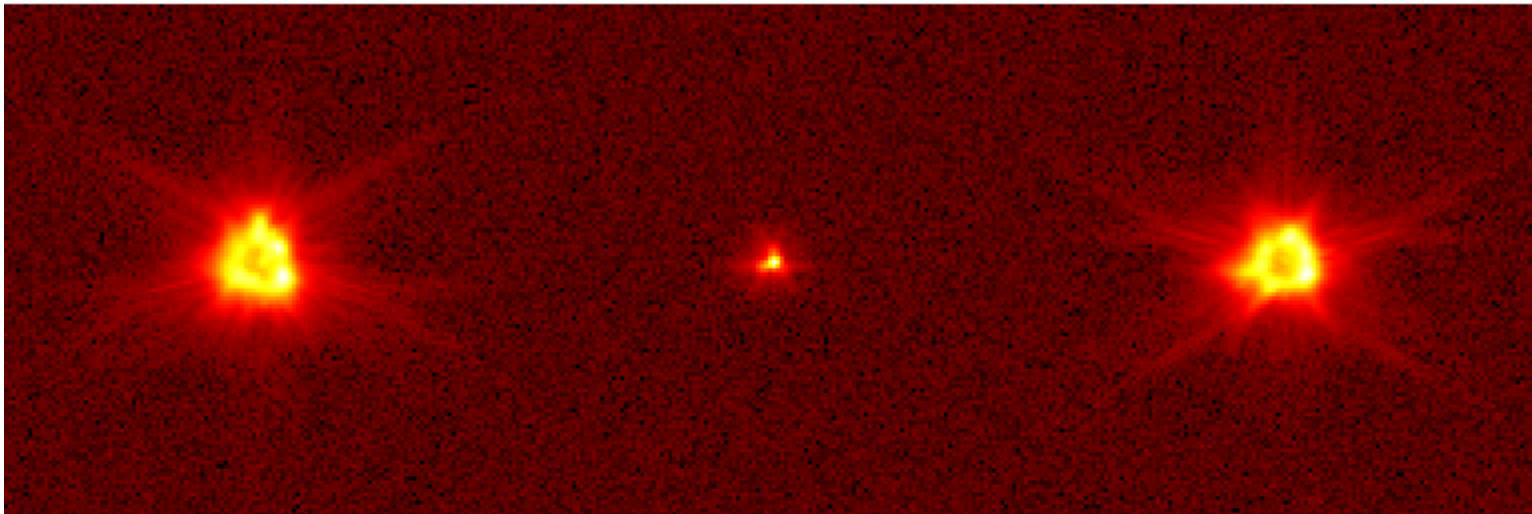
- Monolithic System Results
  - Good results for monolithic system
  - 270x speed improvement
  - Small loss in accuracy
  - How does approximation perform for a segment system?
- Segmented test case
  - Same 1-5  $\mu\text{m}$  FGS-like system
  - Global aberration model: Third order aberrations
  - Segment aberration model: Piston, tip, tilt
  - Monte Carlo simulation
    - 16 trials
    - 0.1 waves RMS wavefront errors

-2 waves defocus

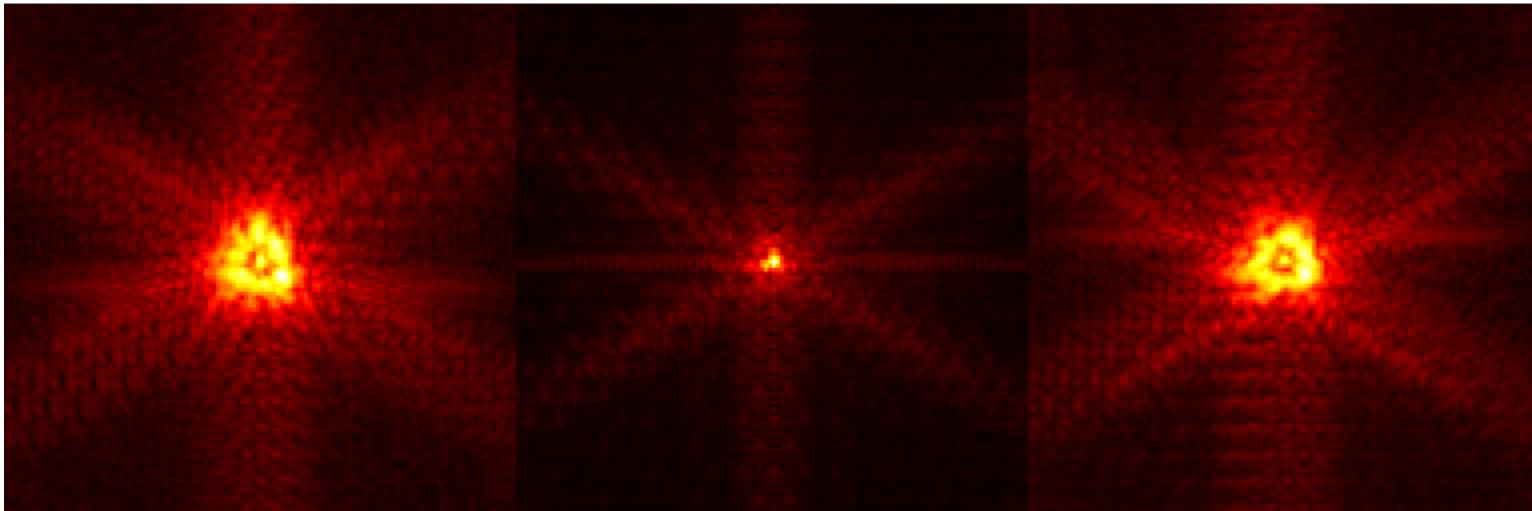
0 waves defocus

+2 waves defocus

Data



Unblurred  
Monochromatic



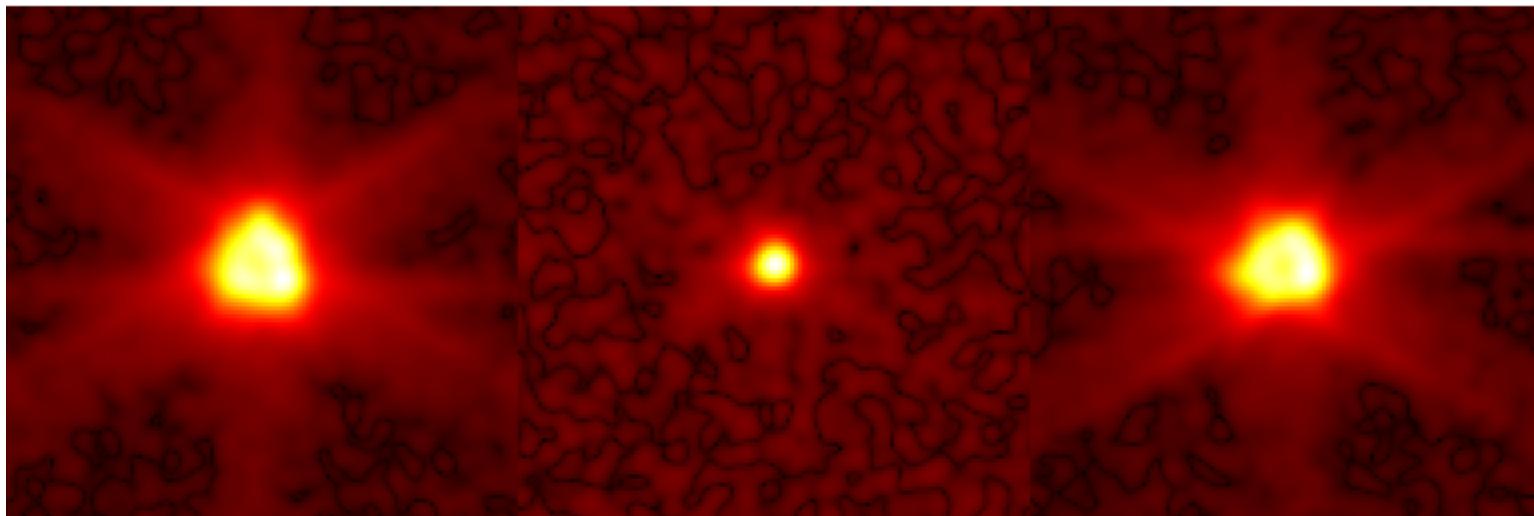
## Blurred PSFs (Segmented)

-2 waves defocus

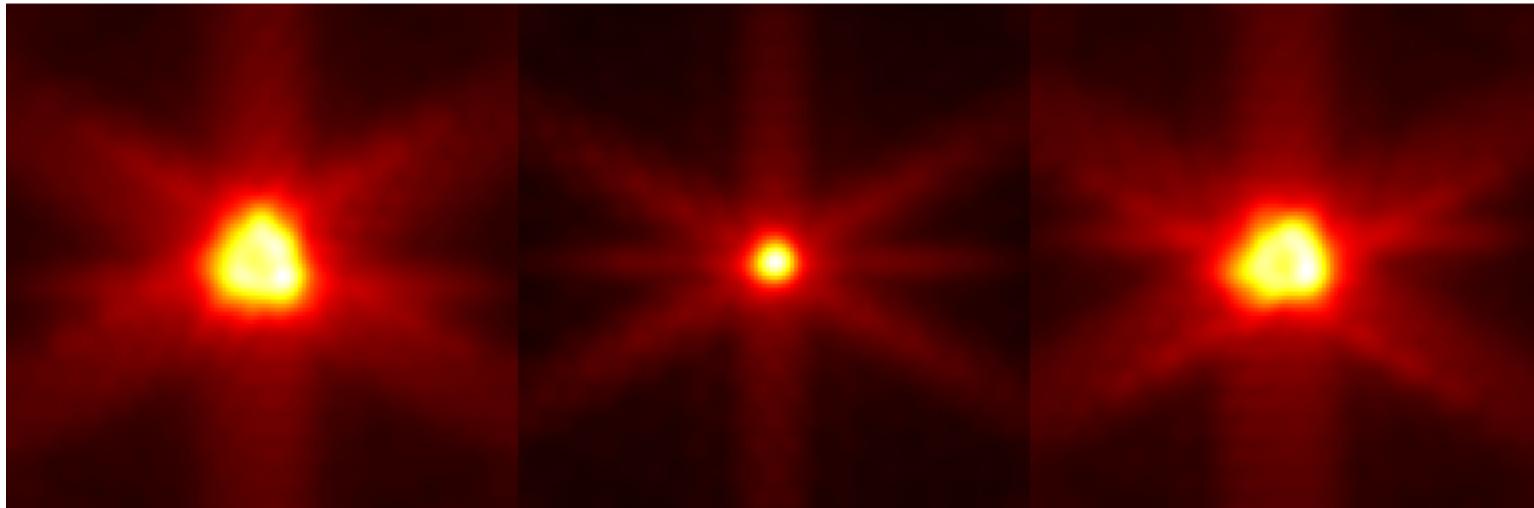
0 waves defocus

+2 waves defocus

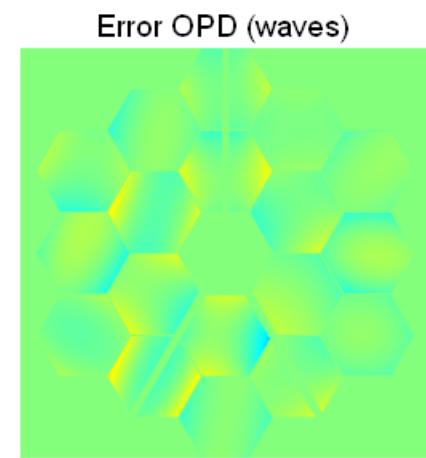
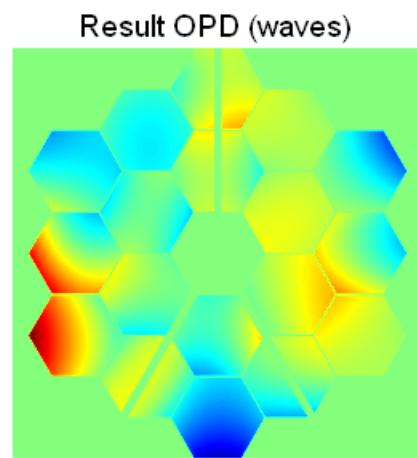
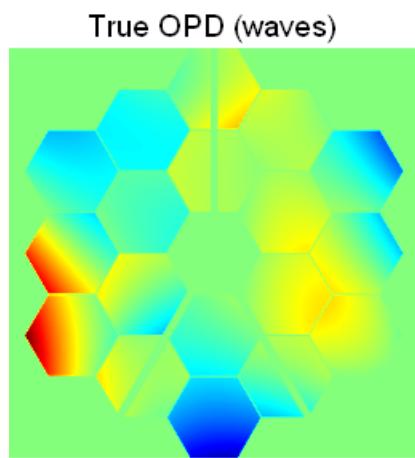
Data

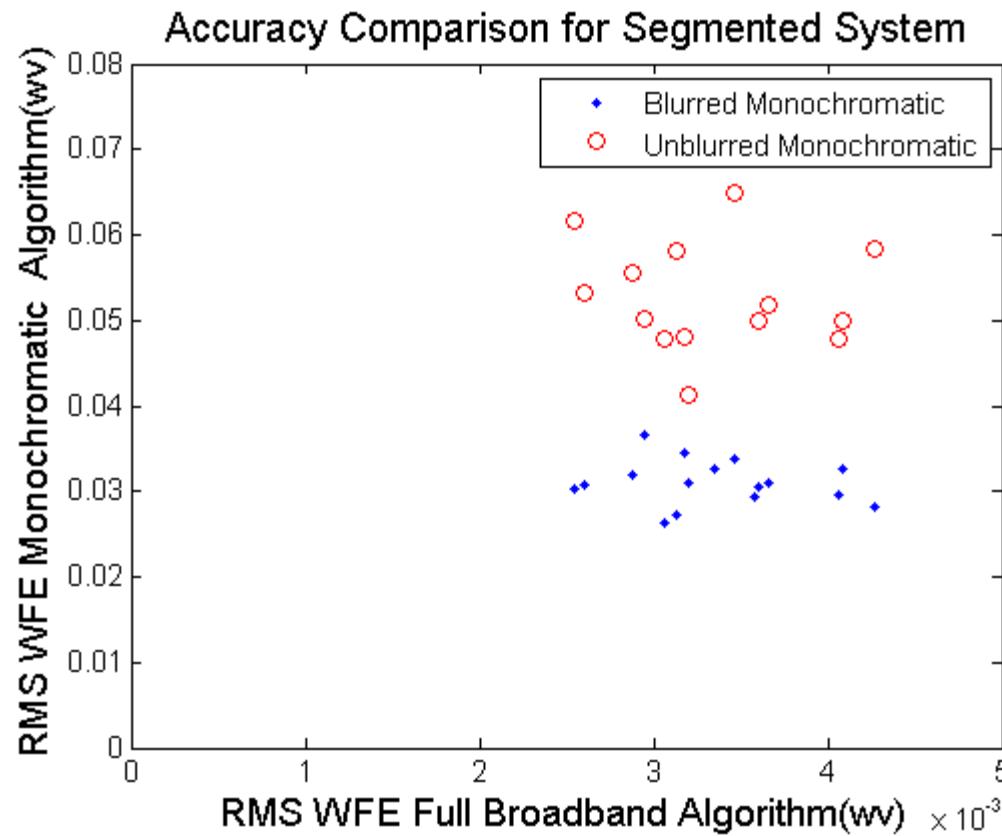


Blurred  
Monochromatic



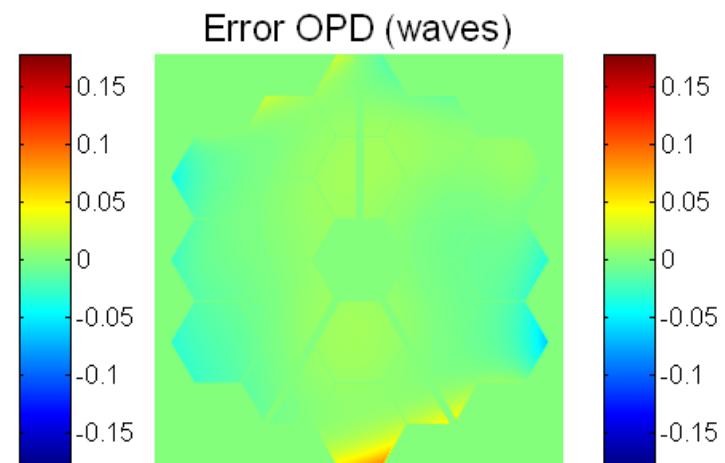
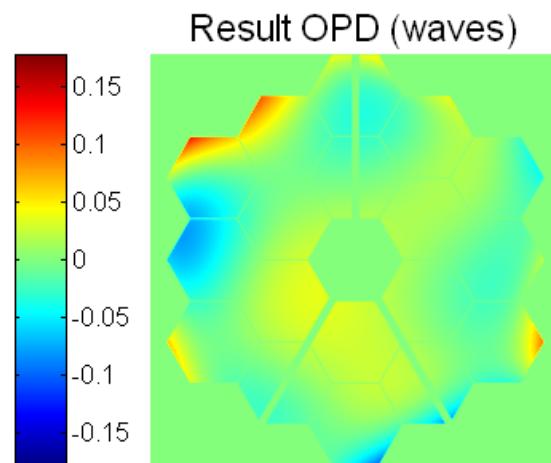
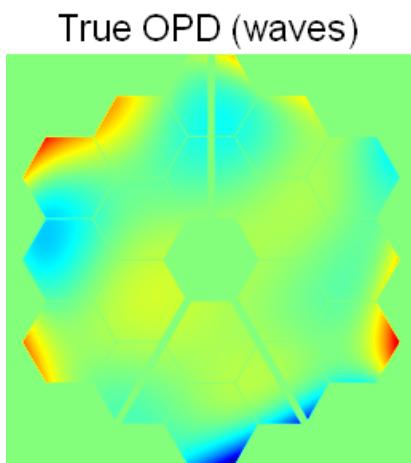
# Results, Segmented System





- Reduced Accuracy with segmented system
- Possible causes
  - More diffraction
  - More variables (higher order phases)
  - Which is responsible?
- To investigate
  - Apply segmented system mask to low order phases
  - Same phase model as monolithic aperture
  - Same aperture mask as segmented system
  - Isolates diffraction effect from higher order phase effect
  - Limit study to 0.25 waves wavefront to avoid issues with large wavefronts

# Results: Segmented Mask

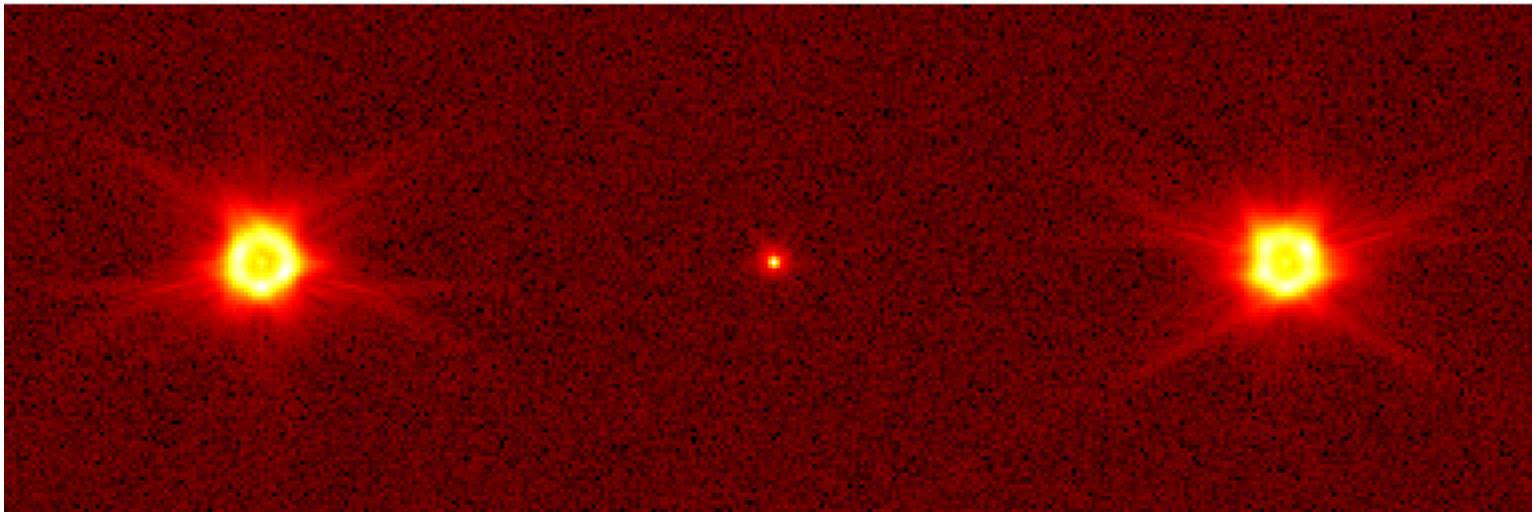


-2 waves defocus

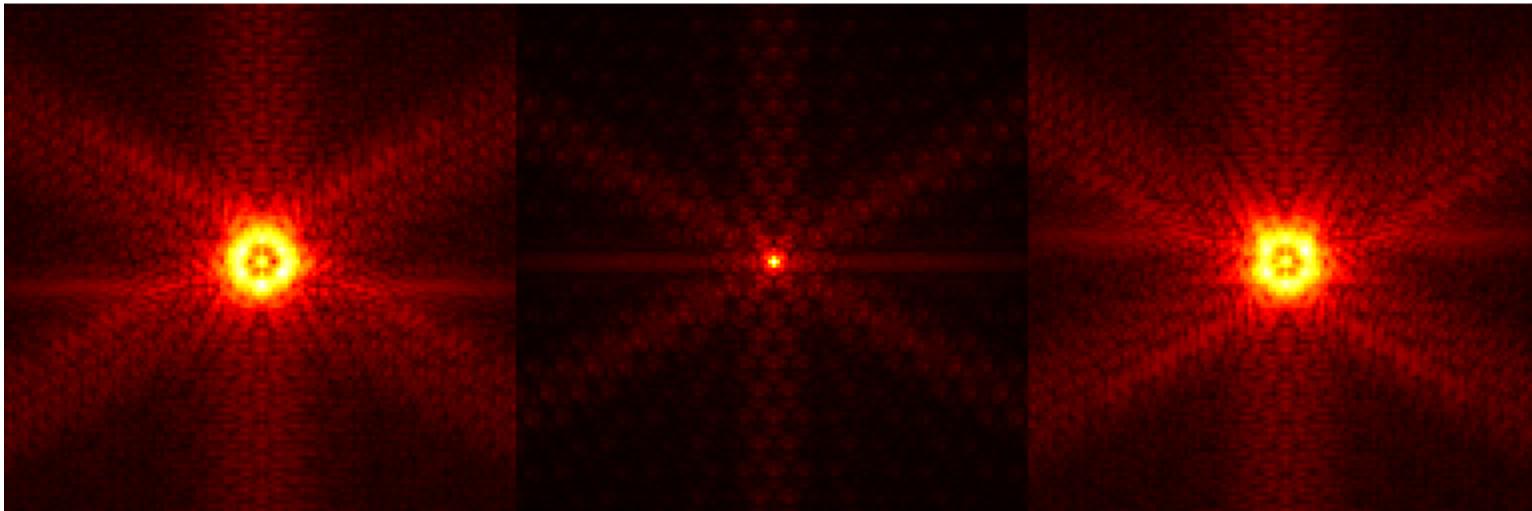
0 waves defocus

+2 waves defocus

Data



Unblurred  
Monochromatic

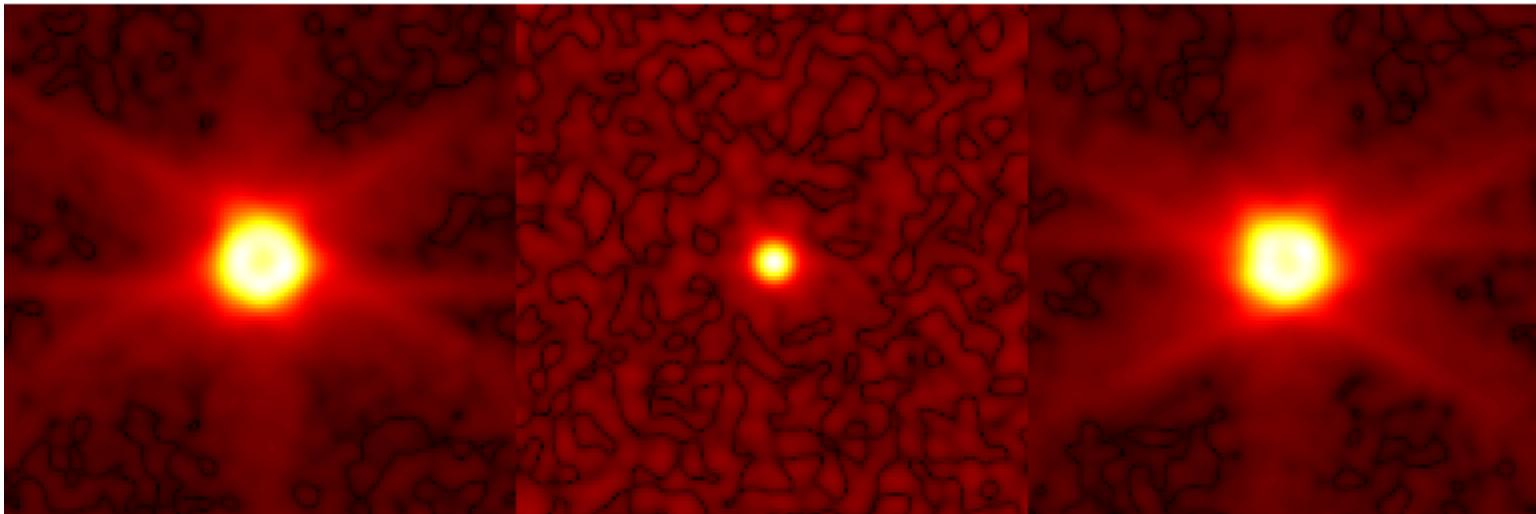


-2 waves defocus

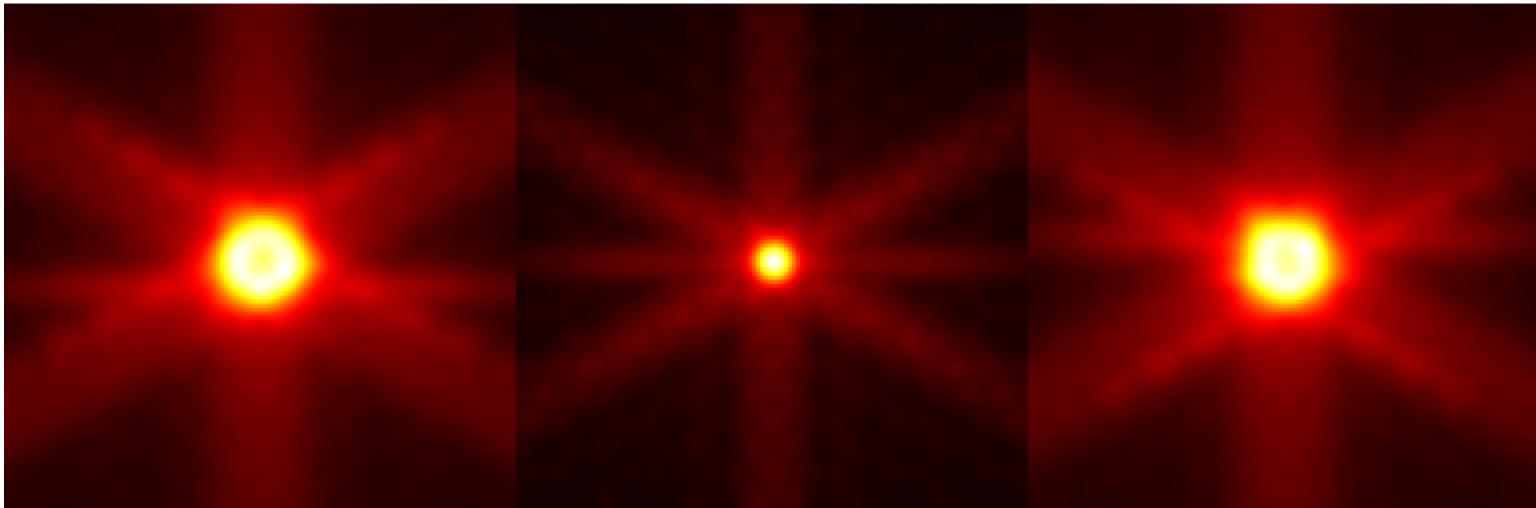
0 waves defocus

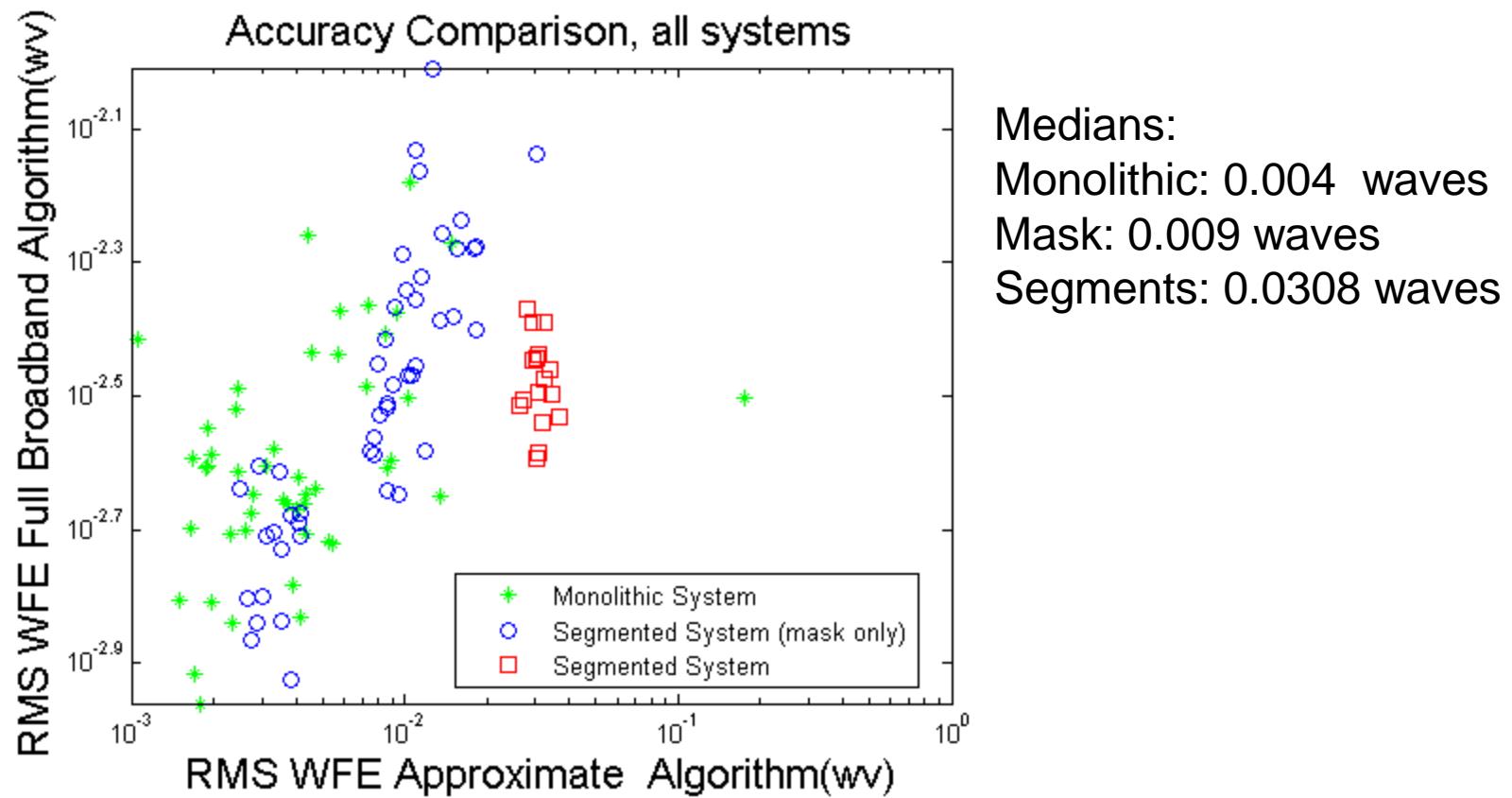
+2 waves defocus

Data



Blurred  
Monochromatic





Both diffraction and higher order phase reduce accuracy!

## Conclusion

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- Broadband phase retrieval needed when:
  - Narrow spectral filters unavailable
  - Dim sources
  - Low throughput due to misalignment
  - Short exposures times
    - Pointing instability (space)
    - Atmospheric instability (ground based AO)
- Traditional approach is computationally burdensome for extreme bandwidths
- Approximate approach
  - Substitute monochromatic model
  - Blur model and data
- Test case performance
  - ~270x reduction in computational cost for FGS-like test case
  - Good accuracy for monolithic system
  - Acceptable accuracy for segmented systems
    - Reduced by diffraction
    - Reduced by higher order segment model

Questions?